

Ensemble Kalman Filter Strategies for Efficient Data Assimilation in Geosciences

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Did you know that...

...your DA performances are widely affected by:

- Sampling
- Tuning



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It's general, it's for everyone, it's for **you!**





Sampling







Sampling (a look to the past)

The **second-order-exact sampling** Pham 1996, Pham 2001

used in SEIK, ETKF and other square root filters

The covariance **P** is approximated by a base **L** and a small symetric matrix **A**: $\mathbf{P} \approx \mathbf{L} \mathbf{A} \mathbf{L}^{\mathsf{T}}$

> The sampling matrix **X** (i.e., the ensemble anomalies) is: **X** = sqrt(EnsSize) **L** S Ω , where **S**² = **A**, $\Omega \Omega^{T} = I$, $\Omega 1 = 0$.

The sampling matches statistical moments up to order 2: **X** $\mathbf{1} = \mathbf{0}$, (1/EnsSize) **X** $\mathbf{X}^{\mathsf{T}} = \mathbf{L} \mathbf{A} \mathbf{L}^{\mathsf{T}}$





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 2nd-order sampling





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- 3 ensemble members:
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- 4 ensemble members:
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- Shady areas represent a Gaussian distribution.
- 3 ensemble members:
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- 4 ensemble members:
 3rd-order sampling
- 4 ensemble members in 3D space: usual 2nd-order sampling



The high-order sampling idea



4 members in 3D (2nd-order approximation)

that project in

4 members in 2D (3rd-order approximation)

that project in

3 weighted members in 1D (5th-order approximation)

Improved precision



by

rising order in the most relevant **PCA** components

NO more members

NO higher computational cost





Enhance your sampling method

The **high-order sampling** Spada et al. 2024 (https://doi.org/10.5194/gmd-2023-170)

used in GHOSH

 $\mathbf{P} \approx \mathbf{L} \mathbf{A} \mathbf{L}^{\mathsf{T}}, \quad \mathbf{S}^2 = \mathbf{A} \\ \mathbf{X} = \mathbf{L} \mathbf{S} \mathbf{E} \mathbf{\Omega}_{\mathsf{h}} \mathbf{W},$

where **W** is the diagonal matrix of the ensemble weights, **S** $L^T L S = E D E^T$ is an eigendecomposition with decreasing eigenvalues, Ω_h is an orthogonal matrix encoding statistical moments.

The sampling matches statistical moments up to an **arbitrary high order** (limited by ensemble size) in the principal error components.



Twin experiment: SEIK vs GHOSH

Toy model: Lorenz96 (62 variables) Observations: odd variables only Observation error: Gaussian noise (standard deviation is 1) Time between observations: 0.1 to 0.3 time units Ensemble Size: 15, 31 and 63 Inflation (forgetting factor): 0.5 to 1.0 Experiments: 400 experiments for each configuration, randomly changing truth, observations and initial conditions, for a total of 66000 tests.



Twin experiment: SEIK vs GHOSH

- GHOSH **always improves RMSE** (up to 70% reduction),
- GHOSH converges for larger intervals between observations (0.25 and 0.3)
- GHOSH is more stable and needs less inflation
- GHOSH has no higher computational cost than SEIK

Very similar results also with the two-scale Lorenz05 model





Realistic 3D test

Setup:

- o Mediterranean Sea
- 1-year simulations
- 1/4° horizontal resolution
- 16 ensemble members
- RMSD to independent data
- 18 tests with different parameters
 (e.g., inflation and sampling order)

Results:

Up to 45% RMSD reduction in a non-assimialted variable (nitrate)



Model (BGC + transport): BFM + OGSTM



Observations: Satellite chlorophyll



Tuning





Tuning what?

Model parameters and initial conditions Filter parameters, e.g., inflation and observation error







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If you have a prior, leave it to filters and sampling methods



Tuning how?

We need an index to optimize. It must be general, data-driven and it should make sense.





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Yes, they are all equivalent!



The auto-tuning minimization

Recall: $\mathbf{P} \approx \mathbf{L} \mathbf{A} \mathbf{L}^{\mathsf{T}}$.

$$\mathbf{P}_{\text{like}} = \mathbf{P}_{\mathbf{H}} + \mathbf{R} = \mathbf{L}_{\mathbf{H}} \mathbf{A} \mathbf{L}_{\mathbf{H}}^{\mathsf{T}} + \mathbf{R},$$

where $\mathbf{L}_{\mathbf{H}}$ is the projection of \mathbf{L} in observation space.

Given that **y** is the observation, \mathbf{y}_{f} is the forecasted observation and $\mathbf{d} = \mathbf{y} - \mathbf{y}_{f}$

$$Loss = |\mathbf{P}_{like}| + \mathbf{d}^T \mathbf{P}_{like}^{-1} \mathbf{d}$$



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It can be computed lightning fast by projecting in ensemble space



Twin experiment #1 (100-tests average)

Auto-tuning:

• forgetting factor

Results:

The filter with auto-tuning (purple) converge faster than the best tuned filter.





Twin experiment #2 (100-tests average)

Auto-tuning:

- forgetting factor and
- observation error

Only the **purple** filter must guess the observation error.

Results:

The filter with auto-tuning (purple) is as good as the best tuned filter.





Auto-tuning 3D implementation

Setup:

- Mediterranean Sea
- 1-year simulations
- 1/24° horizontal resolution
- 24 ensemble members
- 3k cores x 150h =
 450k core hours per run!





Model (BGC + transport): BFM + OGSTM



Observations: Satellite chlorophyll



Take home messages

Sampling:

- the high-order sampling and the GHOSH filter significantly improve performance,
- with near the **same computational cost**.

Tuning:

- the likelihood-based auto-tuning saves time (and money),
- while granting the **best performances**.





THANK YOU!

Take a look at:"GHOSH v1.0.0: a novel Gauss-Hermite High-OrderSampling Hybrid filter for computationally efficientdata assimilation in geosciences"https://doi.org/10.5194/gmd-2023-170

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Instabilities





Long runs





Auto-tuning SEIK and GHOSH

